

Sensitivities of future long baseline experiments in the U.S.

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FNAL/BNL study

- ✓ Joint FNAL & BNL study to compare future long baseline experiments in the US
 - ⇒ see talk by M. Bishai for more details
- ✓ This talk about inputs to sensitivity calculations
- ✓ Two detector techniques considered:
 - × Liquid Argon
 - × Water Cherenkov
- ✓ Two neutrino beams considered:
 - × narrow band NuMI off-axis beam to NOvA location
 - × wide band beam to DUSEL site

Long Baseline Experiments



- ✓ Limit to above experiments for this talk
- ✓ More variations have been studied

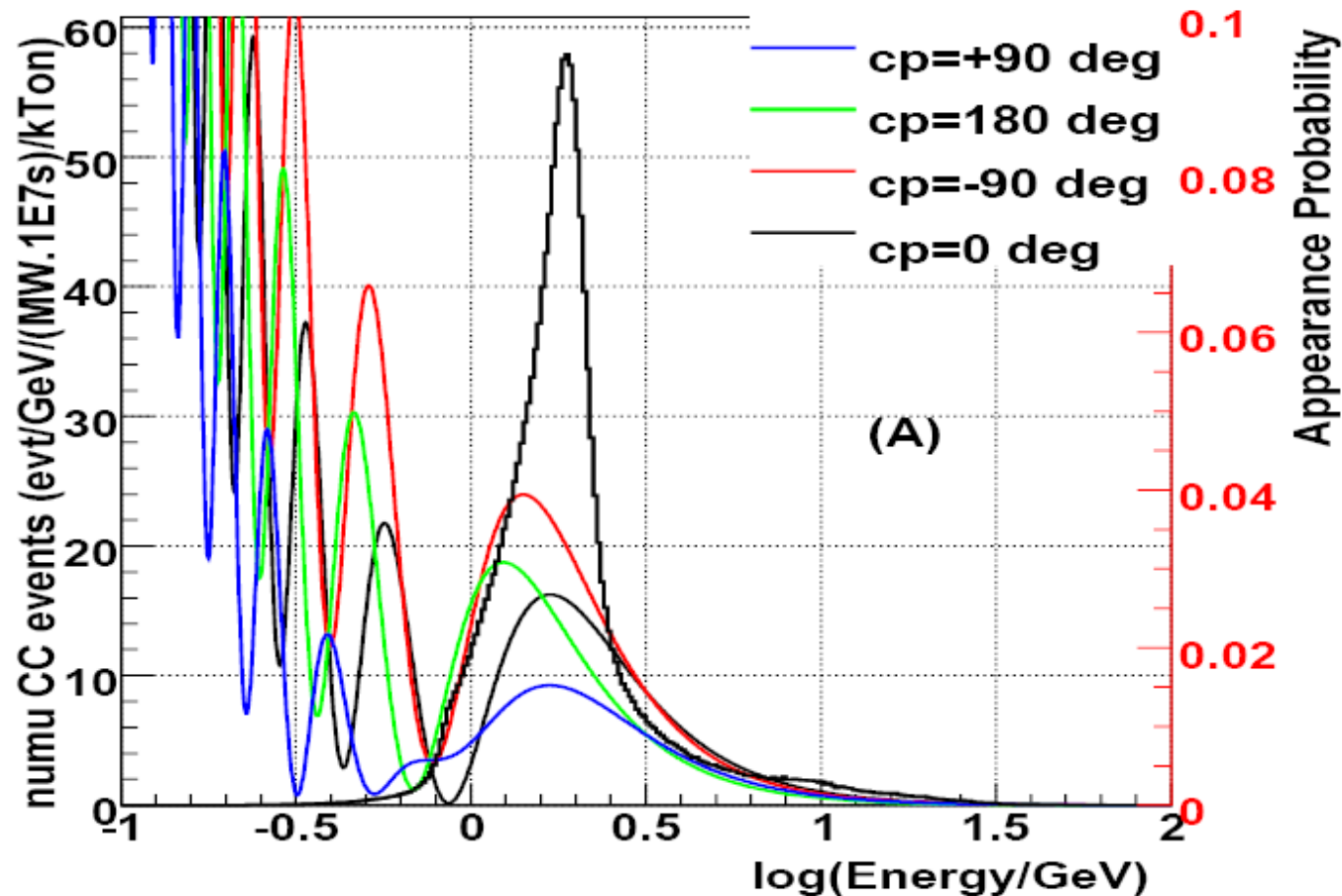
Neutrino Beams

Beam Power

- ✓ Possible scenarios for upgrades to increase beam power output of the Main Injector at FNAL:
⇒ see talk B. Zwaska
- ✓ Calculations for this study mostly done for an exposure of $30 \cdot 10^{20}$ protons on target (pot) for neutrino and anti-neutrino running each.
- ✓ You can plug in your favorite upgrade and convert this number to length of data taking
- ✓ As example:
1.2 MW 120 GeV proton beam
@ FNAL: $1.7 \cdot 10^7$ s/year
→ $10 \cdot 10^{20}$ pot/year

NuMI off-axis

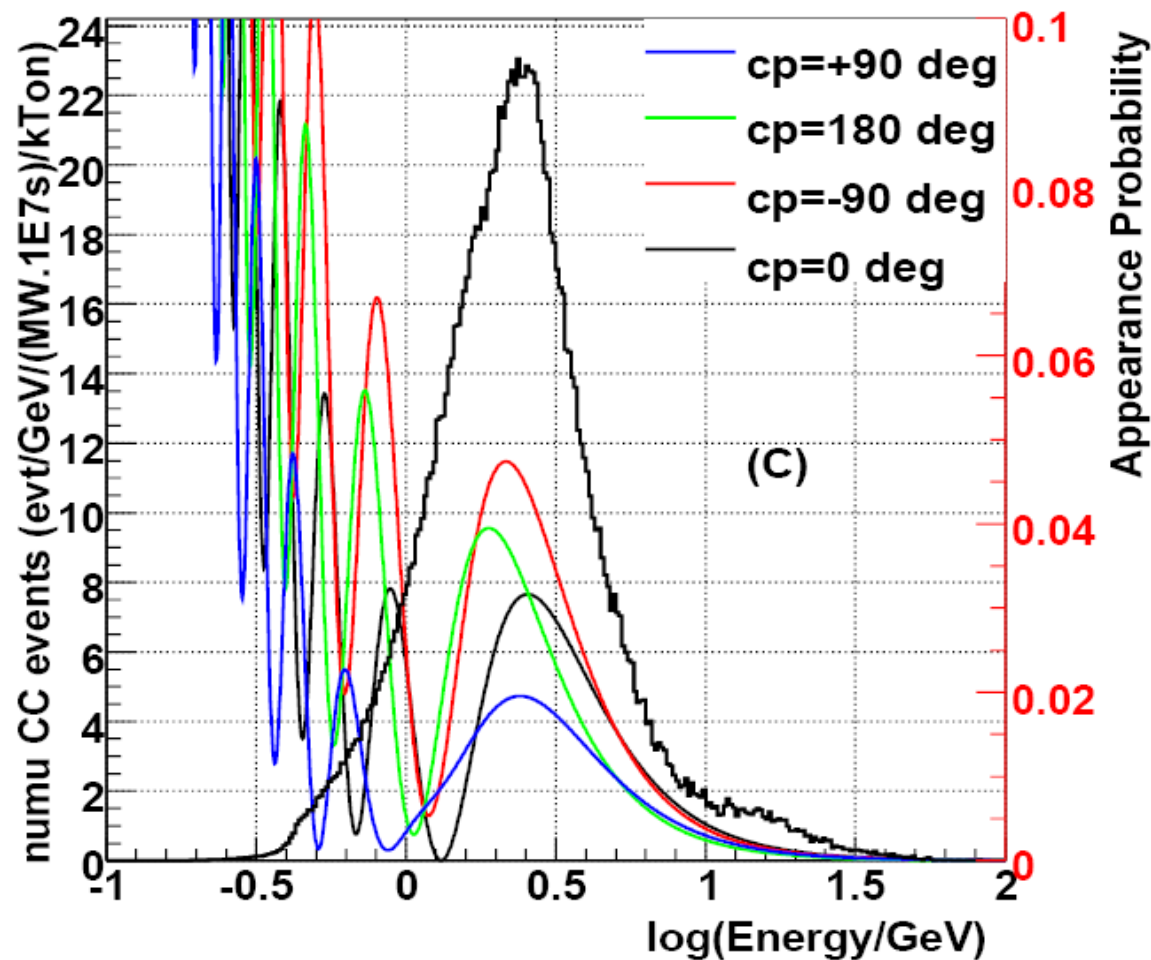
- ✓ ν_μ cc event rates and ν_e appearance probabilities for a 0.8° off-axis NuMI beam at 810km



- ✓ sensitivities for other options, like placing a detector at 2nd oscillation maximum considered

Wideband beam

- ✓ ν_μ cc event rates and ν_e appearance probabilities for a 120 GeV 0.5° off-axis wide band beam at 1300km



- ✓ Other proton energies and distances also considered

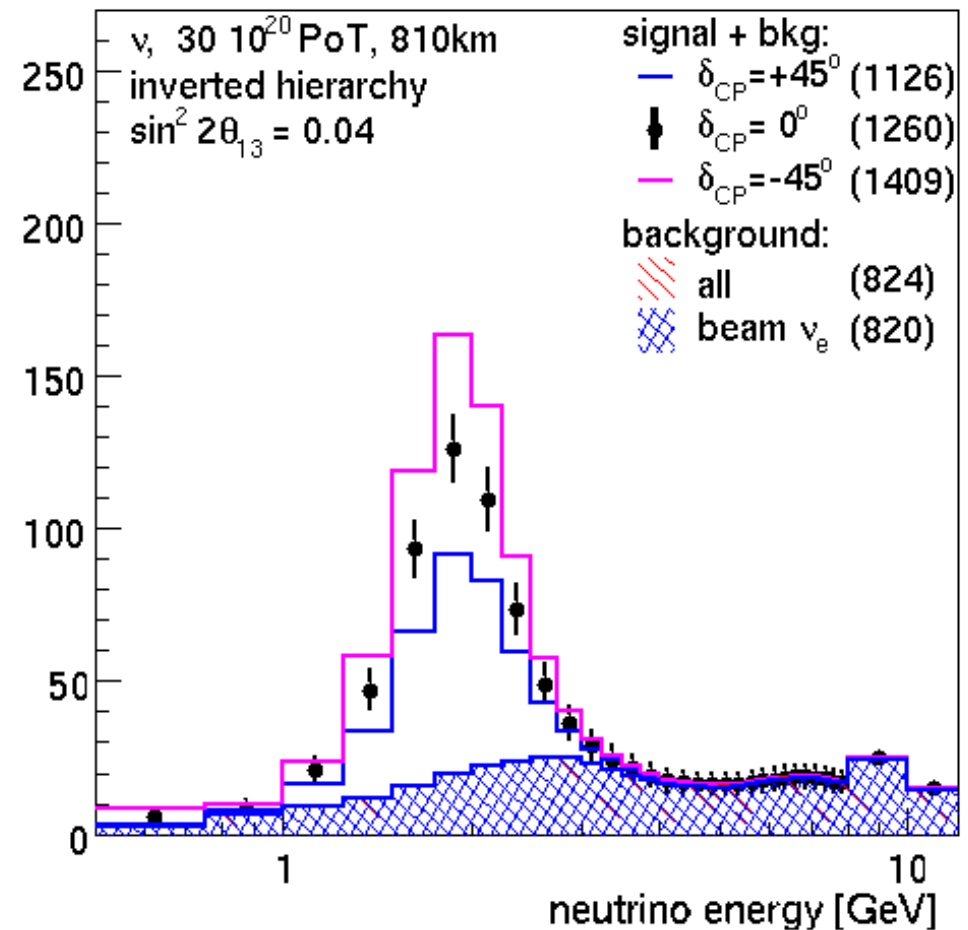
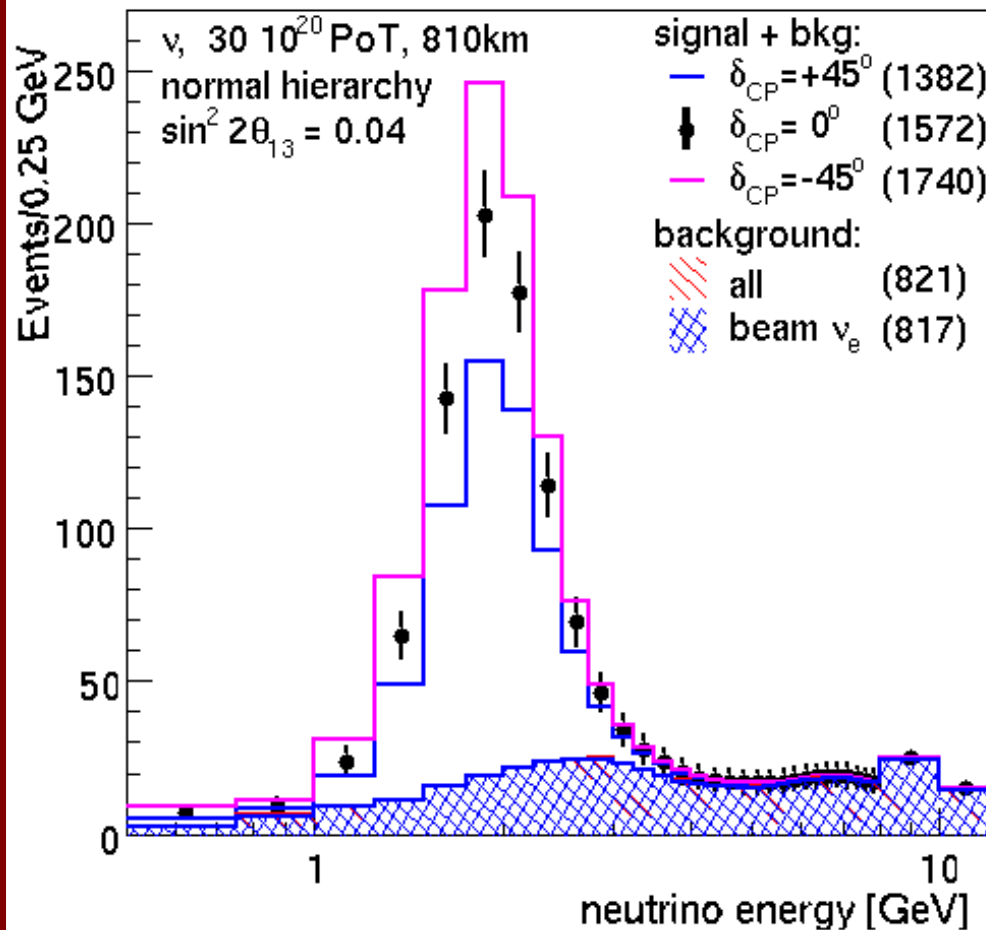
Detectors

Liquid Argon

- ✓ 100 kt liquid argon (LAr) detector
- ✓ Can be placed at NOvA or DUSEL site
- ✓ Signal & background efficiencies based on hand scanning. (B. Flemming et al):
 - ν_e CC efficiency: 80%
 - complete rejection of NC background
- ✓ Results confirmed/exceeded using initial automated tools. (A Curioni)
- ✓ Energy resolution for
 - QE CC: 5% \sqrt{E}
 - non-QE CC: 20% \sqrt{E}

LAr spectra @ NOvA

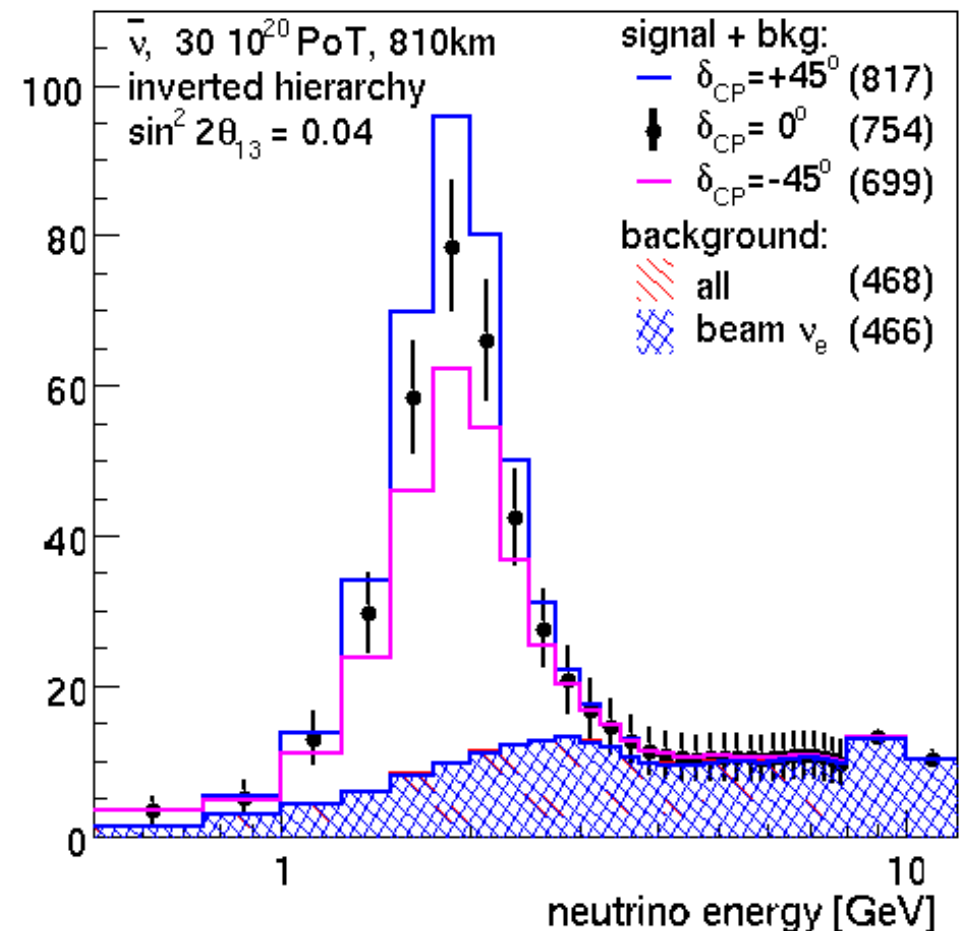
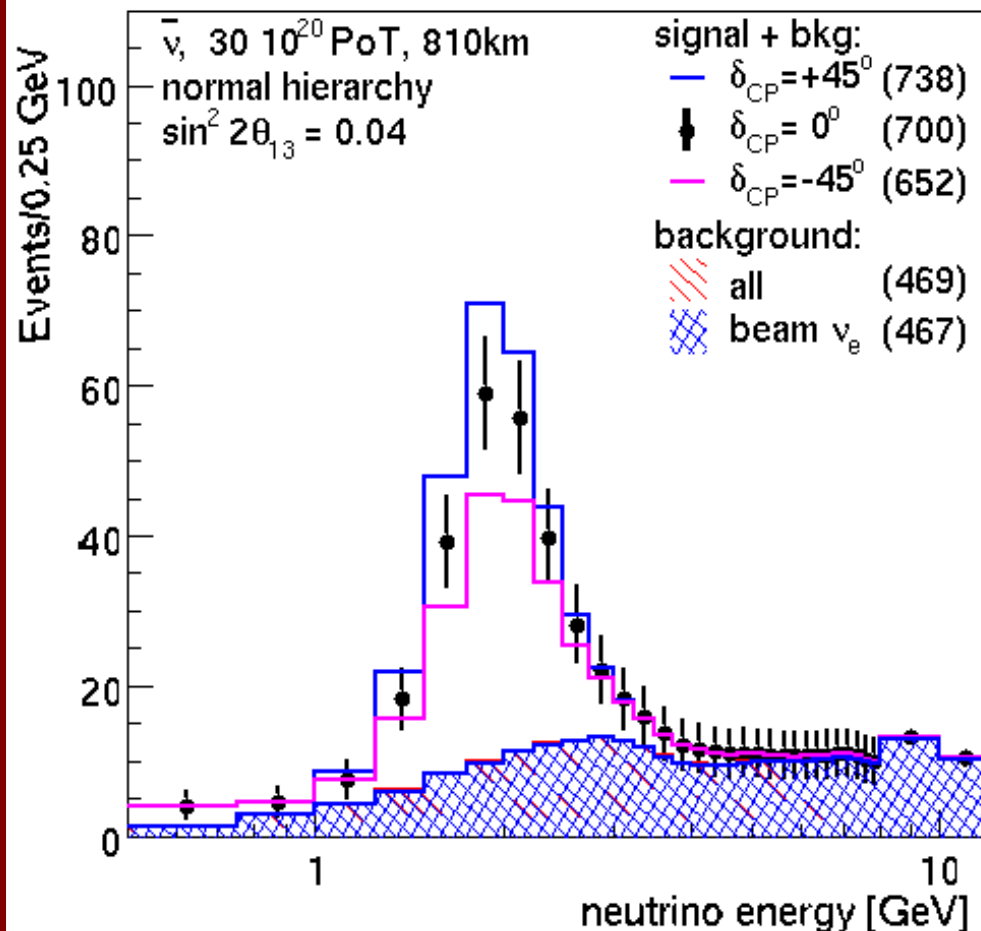
✓ Neutrino running and $\sin^2 2\theta_{13} = 0.04$



put S/B (<4GeV)

LAr spectra @ NOvA

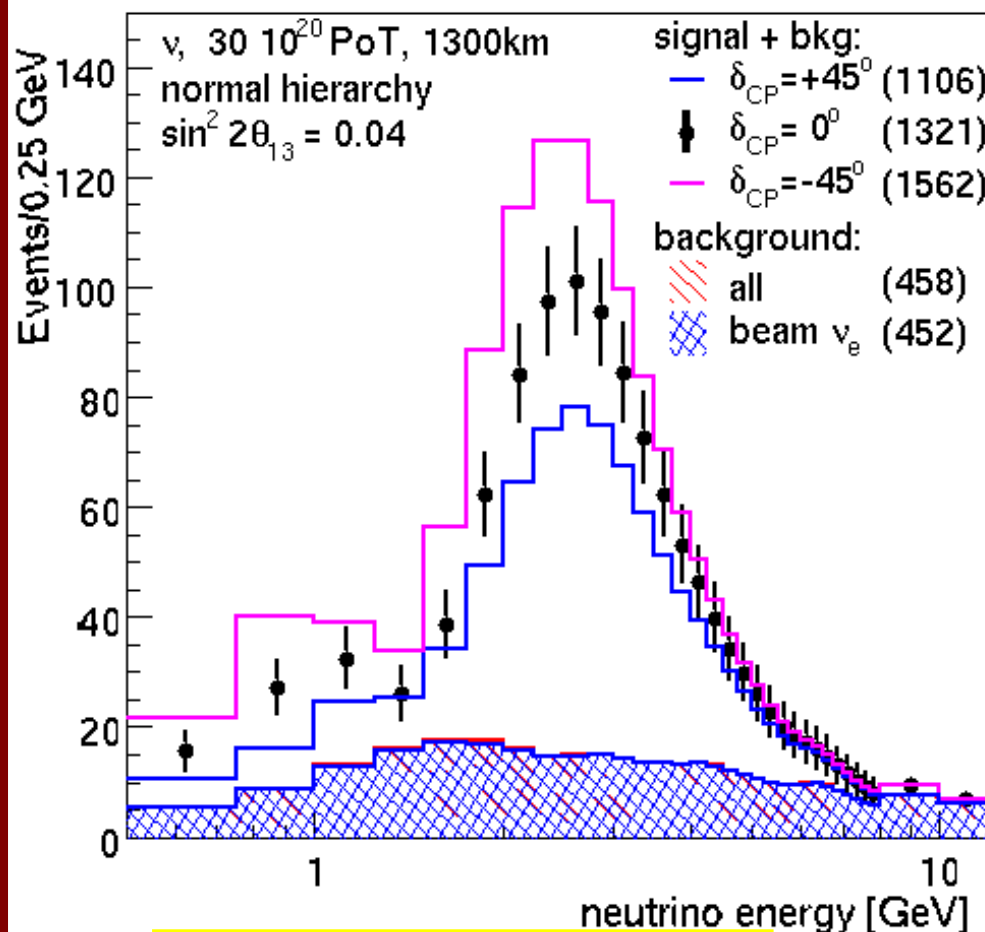
✓ Anti-neutrino running and $\sin^2 2\theta_{13} = 0.04$



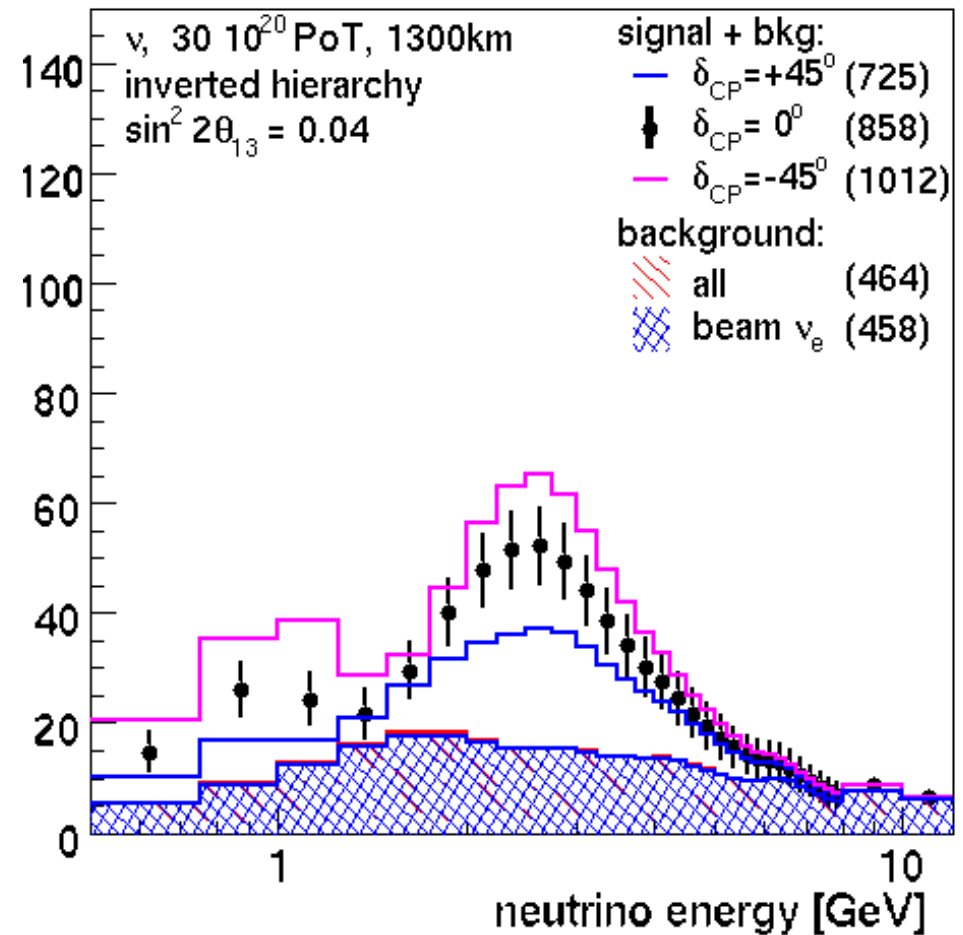
put S/B (<4GeV)

LAr spectra @ DUSEL

✓ Neutrino running and $\sin^2 2\theta_{13} = 0.04$

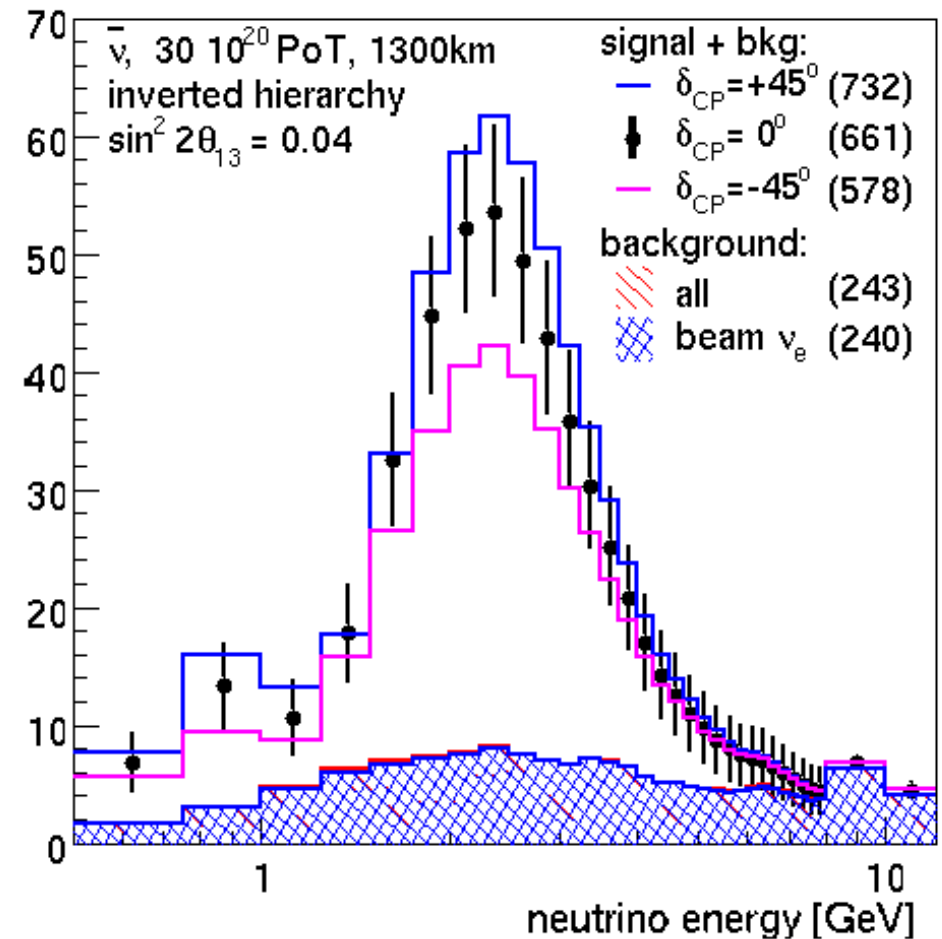
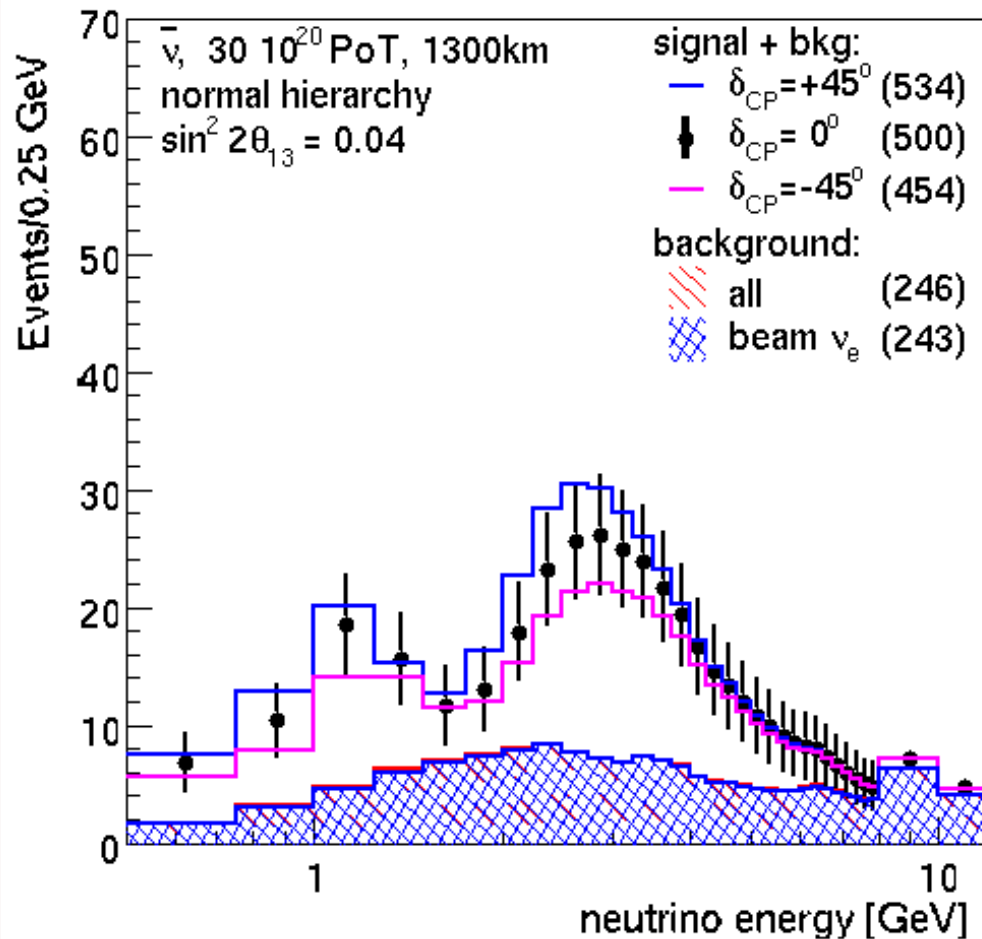


put S/B (<8GeV)



LAr spectra @ DUSEL

✓ Anti-neutrino running and $\sin^2 2\theta_{13} = 0.04$



put S/B (<8GeV)

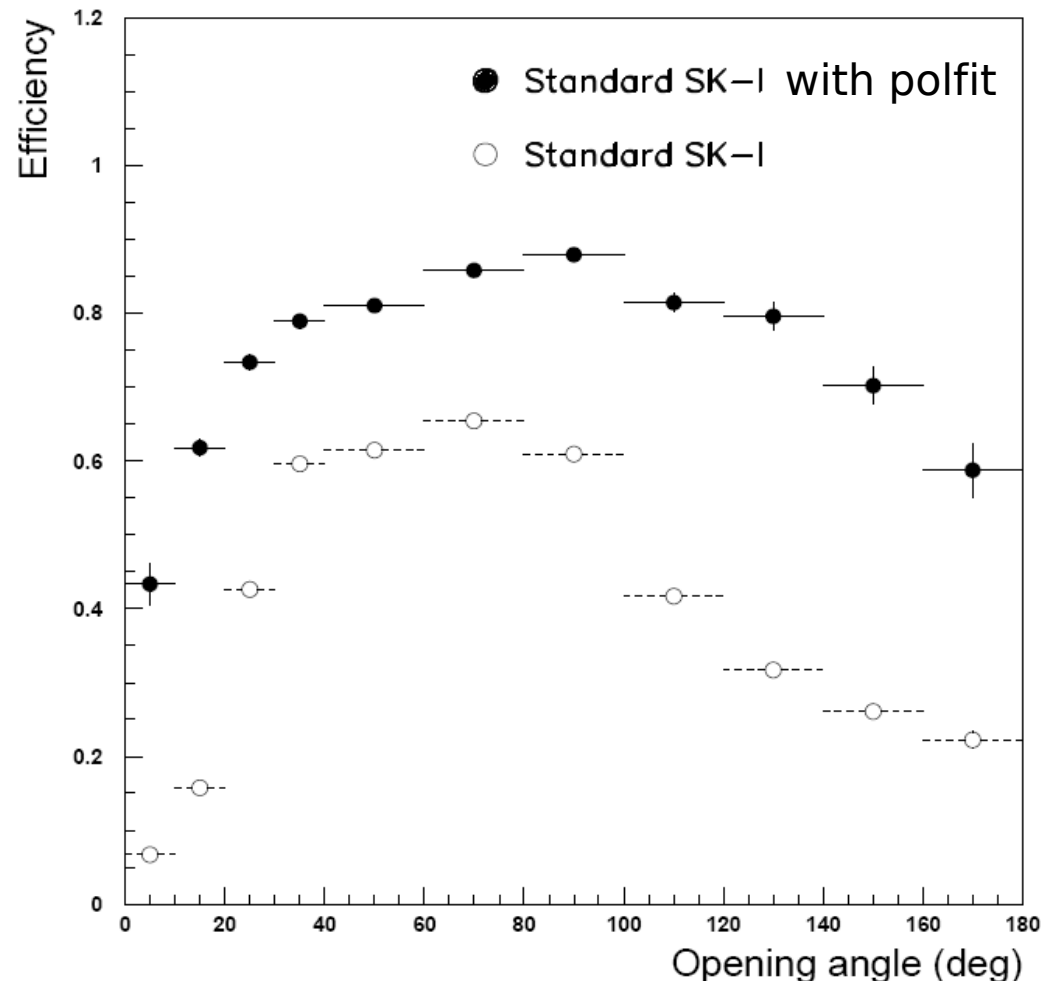
Water Cherenkov Detector

- ✓ 300 kton water Cherenkov (WCh) detector
- ✓ Must be underground to reduce cosmics rate
-> DUSEL
- ✓ Two independent studies performed to improve selection of electron neutrino interactions
C. Yanagisawa
F. Dufour/E. Kearns
- ✓ Both use SuperK MC and follow similar strategy to improve separation between electrons and π^0 decays: very similar results are obtained

Water Cherenkov Detector

- ✓ Standard SuperK cuts are used to select single e-like events

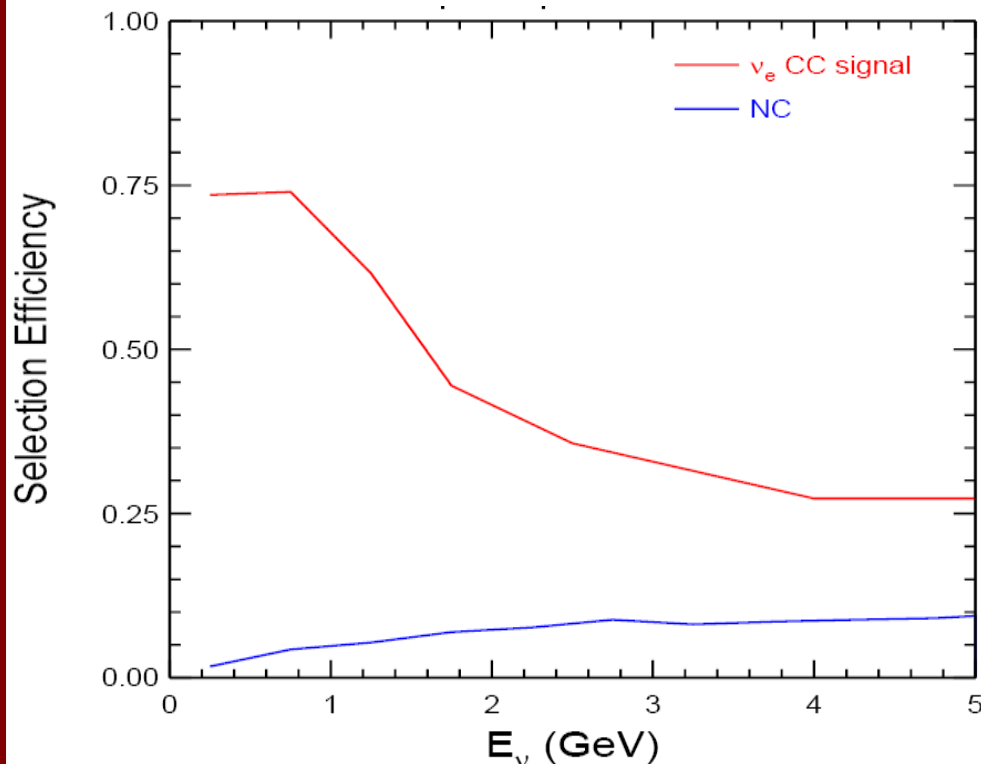
- ✓ A π^0 reconstruction algorithm called Pattern-of-light fit (Polfit) is used to find second ring: improves π^0 eff. by 20-30%



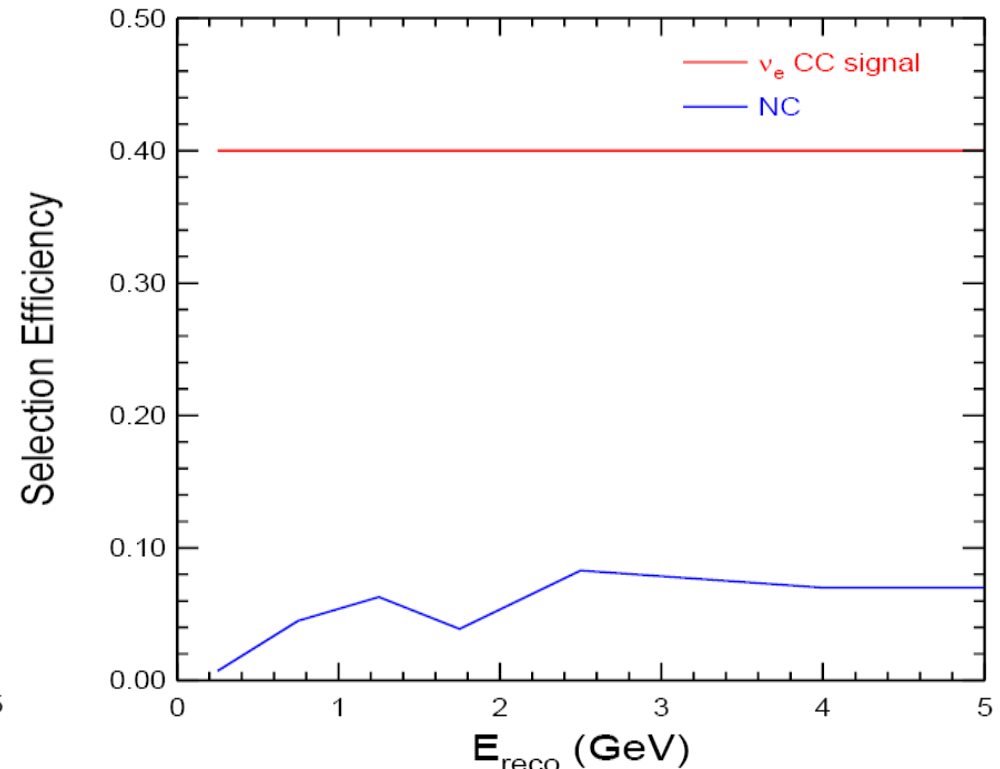
Water Cherenkov Detector

- ✓ The output of polfit and several other variables (some related to knowledge of beam direction) are used as input for a likelihood based analysis

SuperK pre-selection



Likelihood Analysis



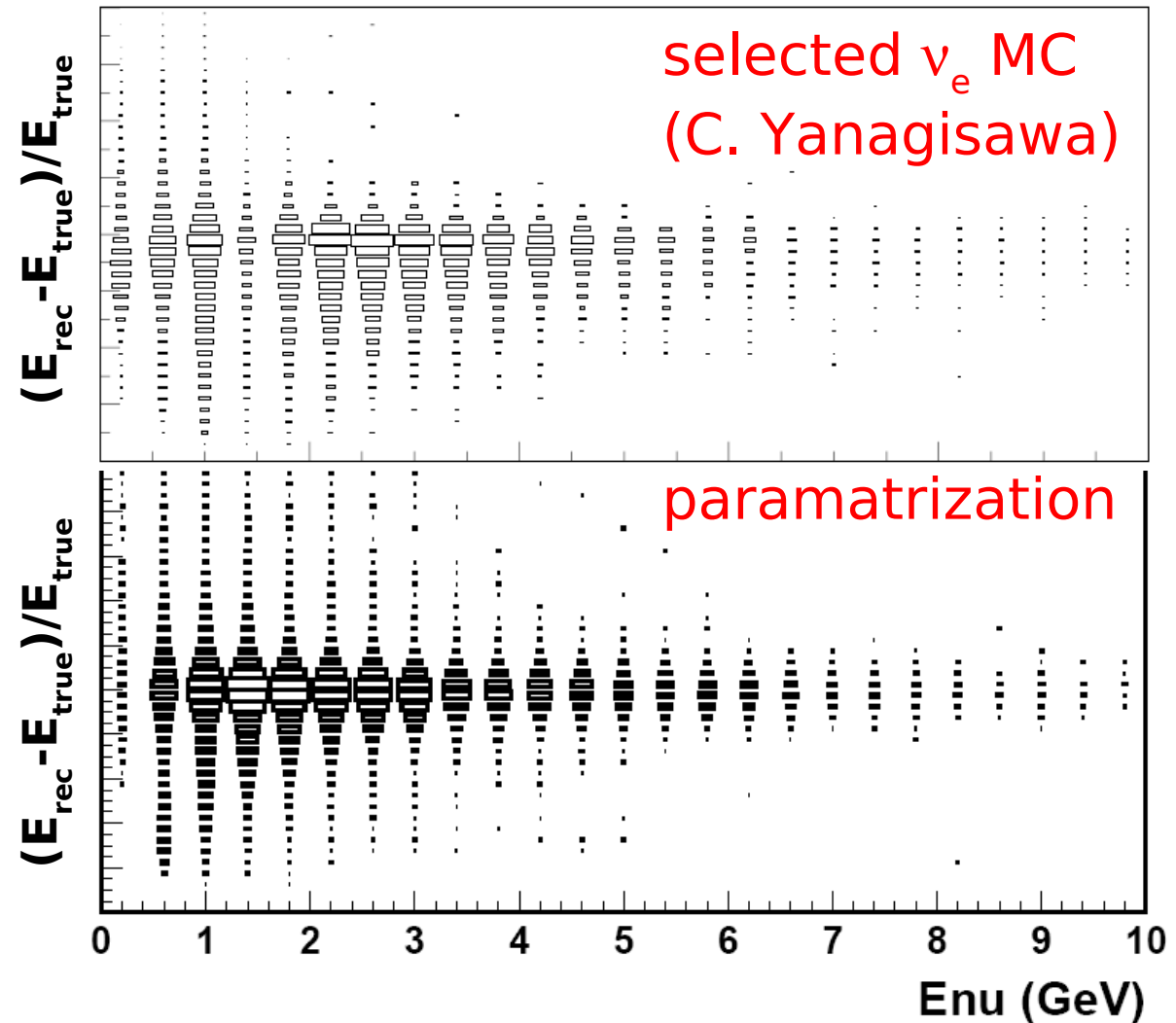
- ✓ Efficiencies are used for sensitivity calculations

Water Cherenkov Detector

✓ Parametrized smearing functions

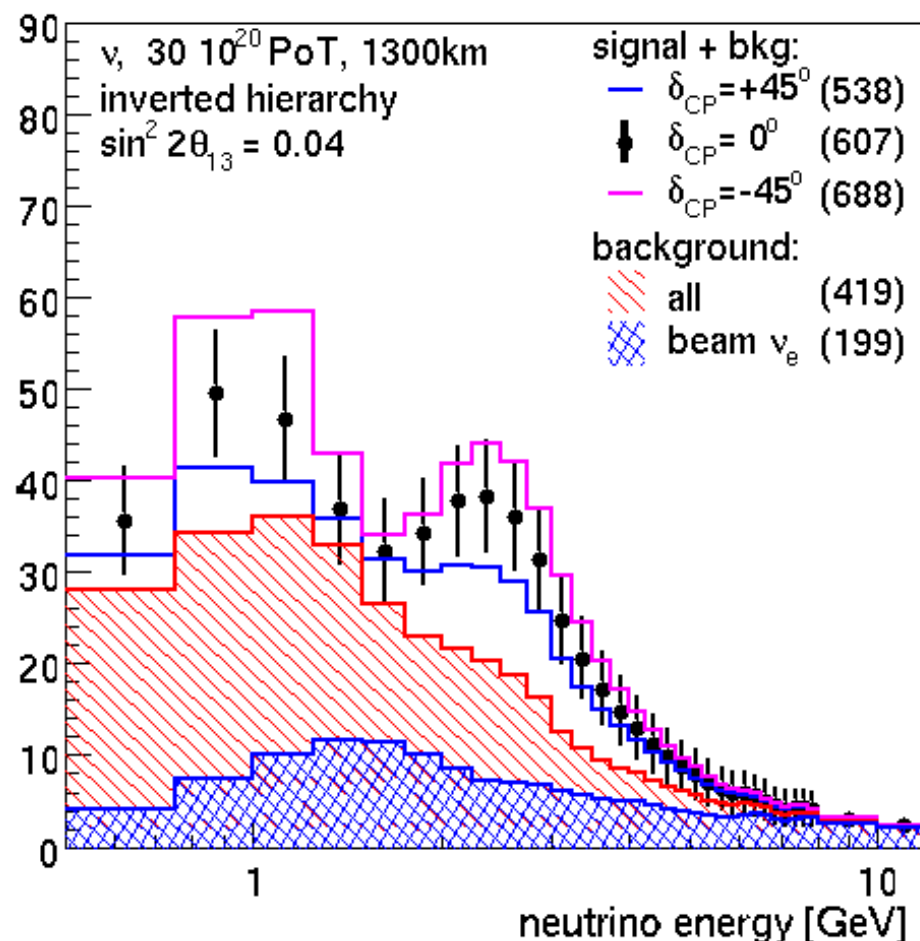
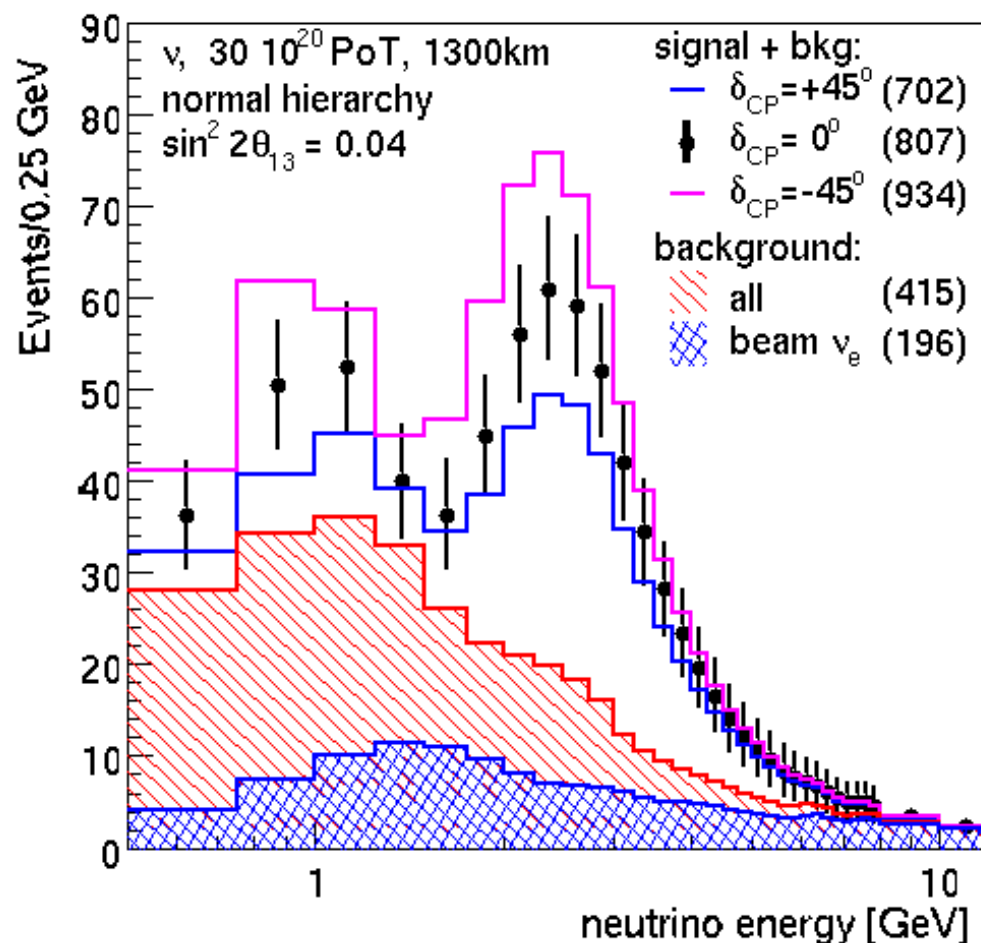
✓ ν_e CC: mostly QE
 $\sigma(E) \sim 10\%$
at 1 GeV

✓ ν_x NC: based
on nuance



WCh spectra @ DUSEL

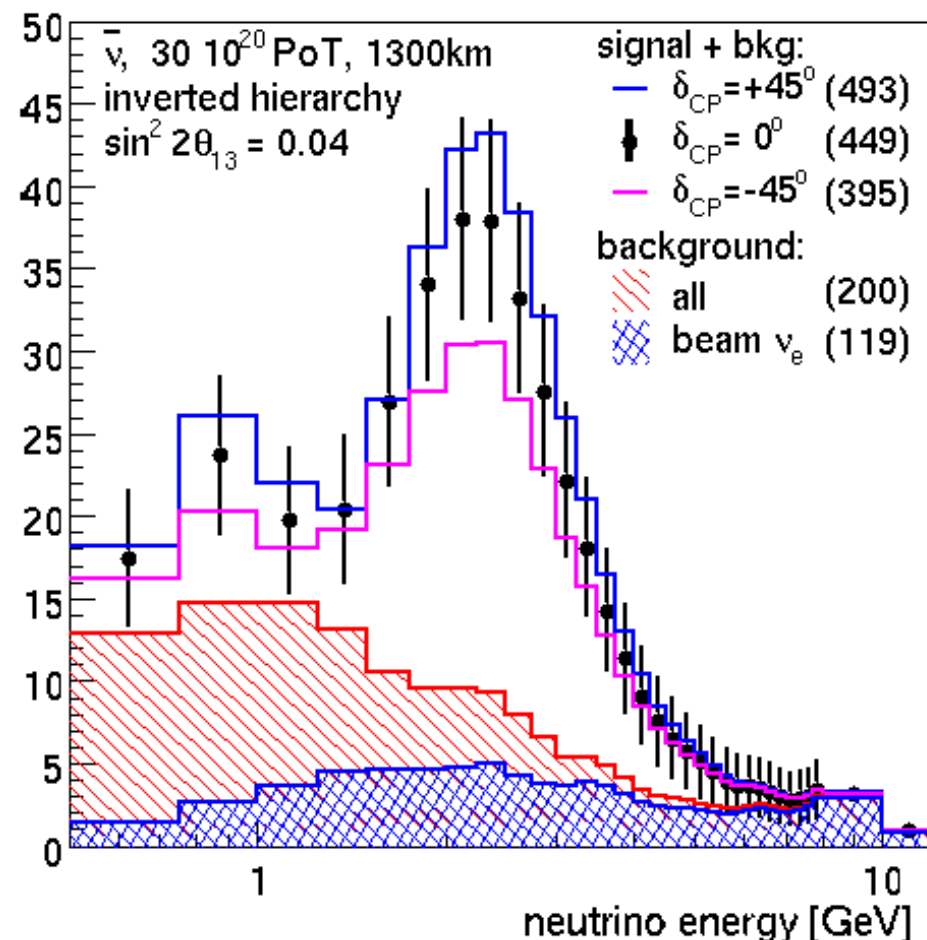
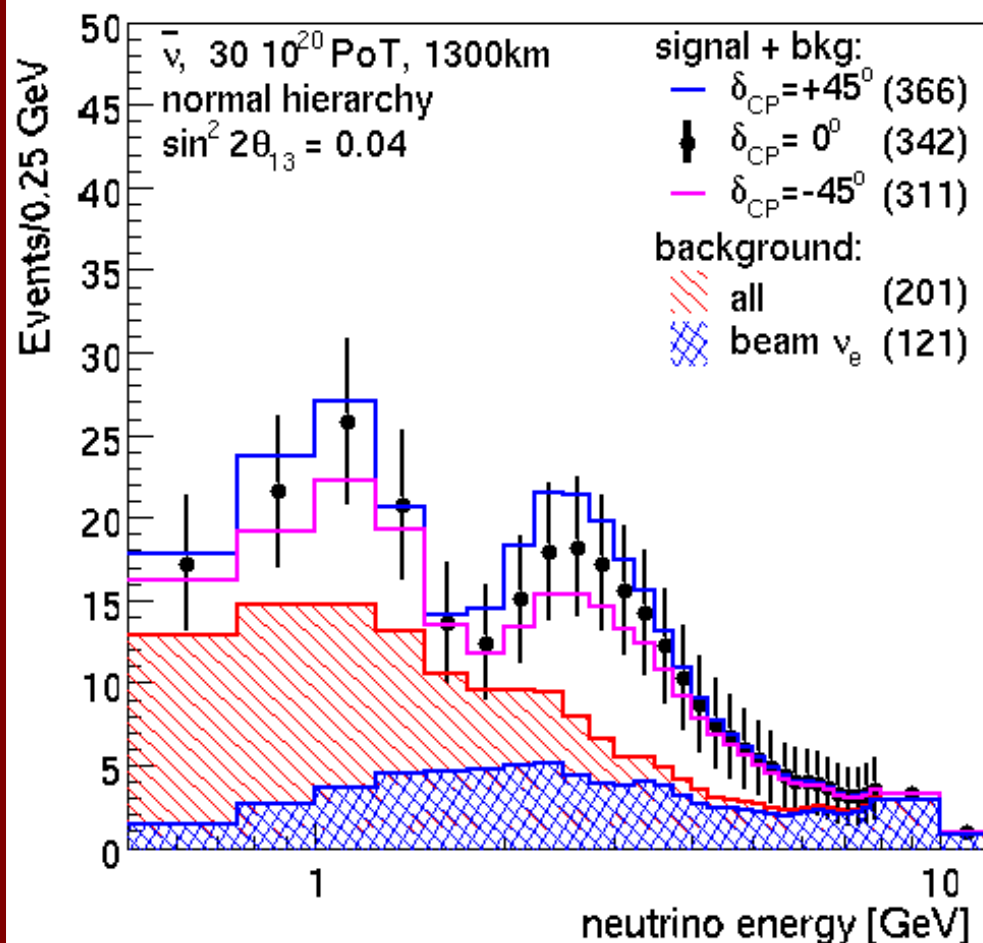
✓ Neutrino running and $\sin^2 2\theta_{13} = 0.04$



put S/B (<8GeV)

WCh spectra @ DUSEL

✓ Anti-neutrino running and $\sin^2 2\theta_{13} = 0.04$



put S/B (<8GeV)

Sensitivity Studies

Sensitivity calculations

- ✓ **Sensitivity to non-zero θ_{13} :** fit spectrum generated for particular $(\theta_{13}, \delta_{cp})$ to hypothesis with $\theta_{13}=0$
- ✓ **Sensitivity to CP violation:** fit spectrum generated for particular $(\theta_{13}, \delta_{cp})$ to hypotheses $\delta_{cp}=0$ and π . Take worst χ^2 . θ_{13} is allowed to float in fit
- ✓ **Sensitivity to matter hierarchy:** fit spectrum for particular $(\theta_{13}, \delta_{cp})$ to hypothesis with opposite mass hierarchy. Both θ_{13} and δ_{cp} are allowed to float
- ✓ **$(\theta_{13}, \delta_{cp})$ measurement (for DUSEL only):** parameter measurement for certain values

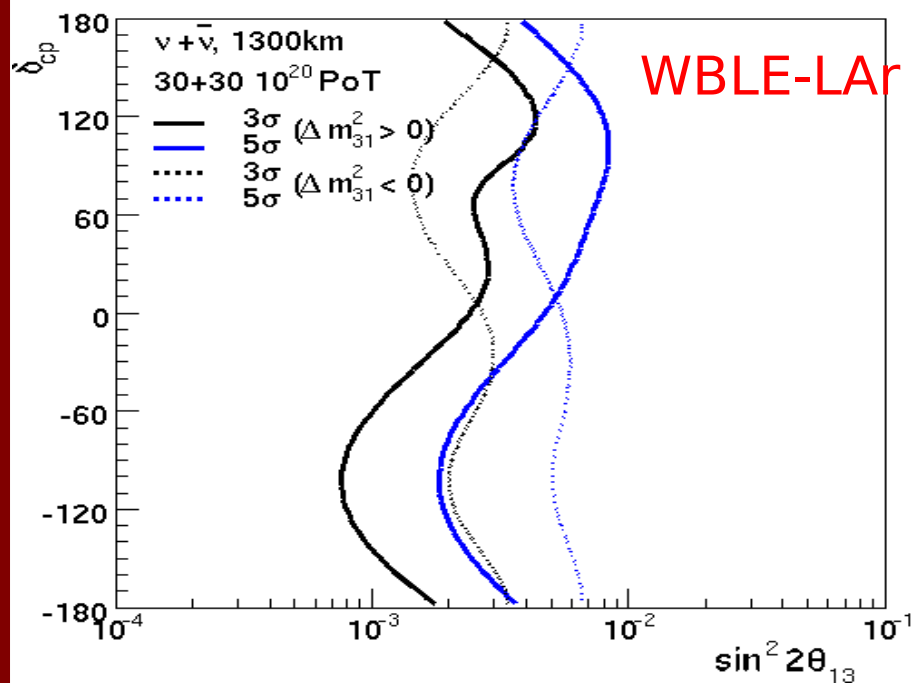
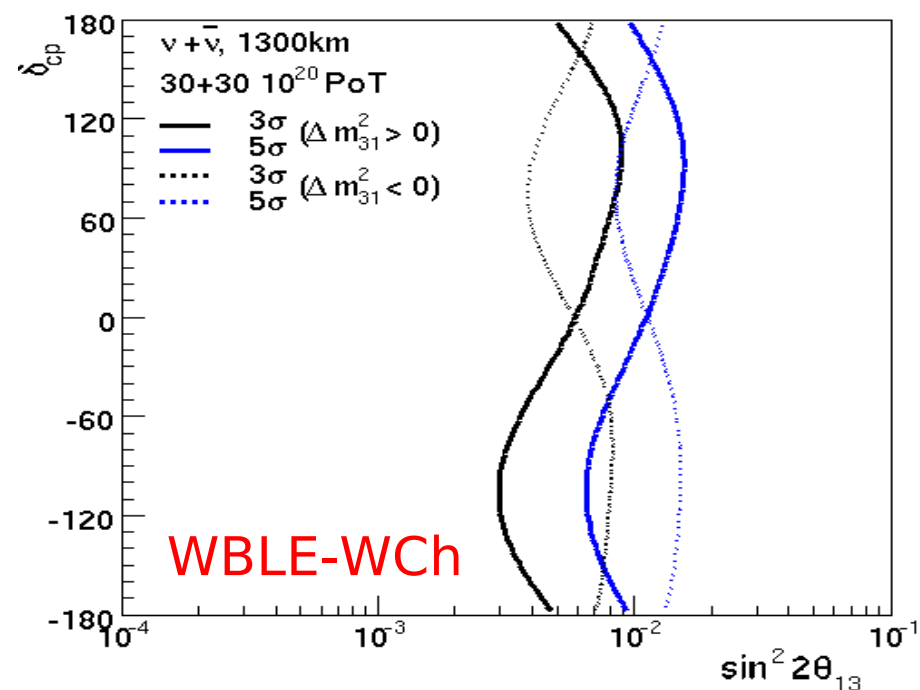
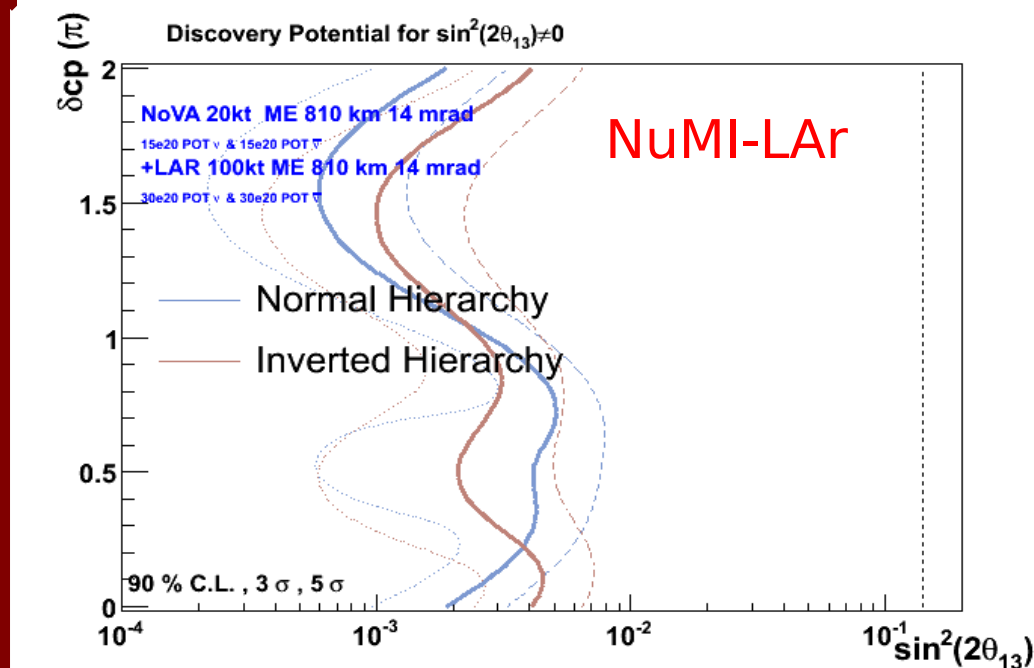
Input parameters & uncertainties

parameter	Uncertainty	
	NuMI off axis	FNAL-DUSEL
$\Delta m^2_{32} = 2.7 \cdot 10^{-3} \text{ eV}^2$	fixed	from $\nu_\mu \rightarrow \nu_\mu$
$\sin^2 \theta_{23} = 1.0$	fixed	from $\nu_\mu \rightarrow \nu_\mu$
$\Delta m^2_{21} = 8.6 \cdot 10^{-5} \text{ eV}^2$	fixed	5%
$\sin^2 2\theta_{12} = 0.86$	fixed	5%
$\rho = ???$	fixed	5%
background	5%	10%
tools	custom program (N. Saoulidou)	GLOBES (M. Dierckxsens, M. Bishai, M. Diwan,...)

Sensitivity calculations

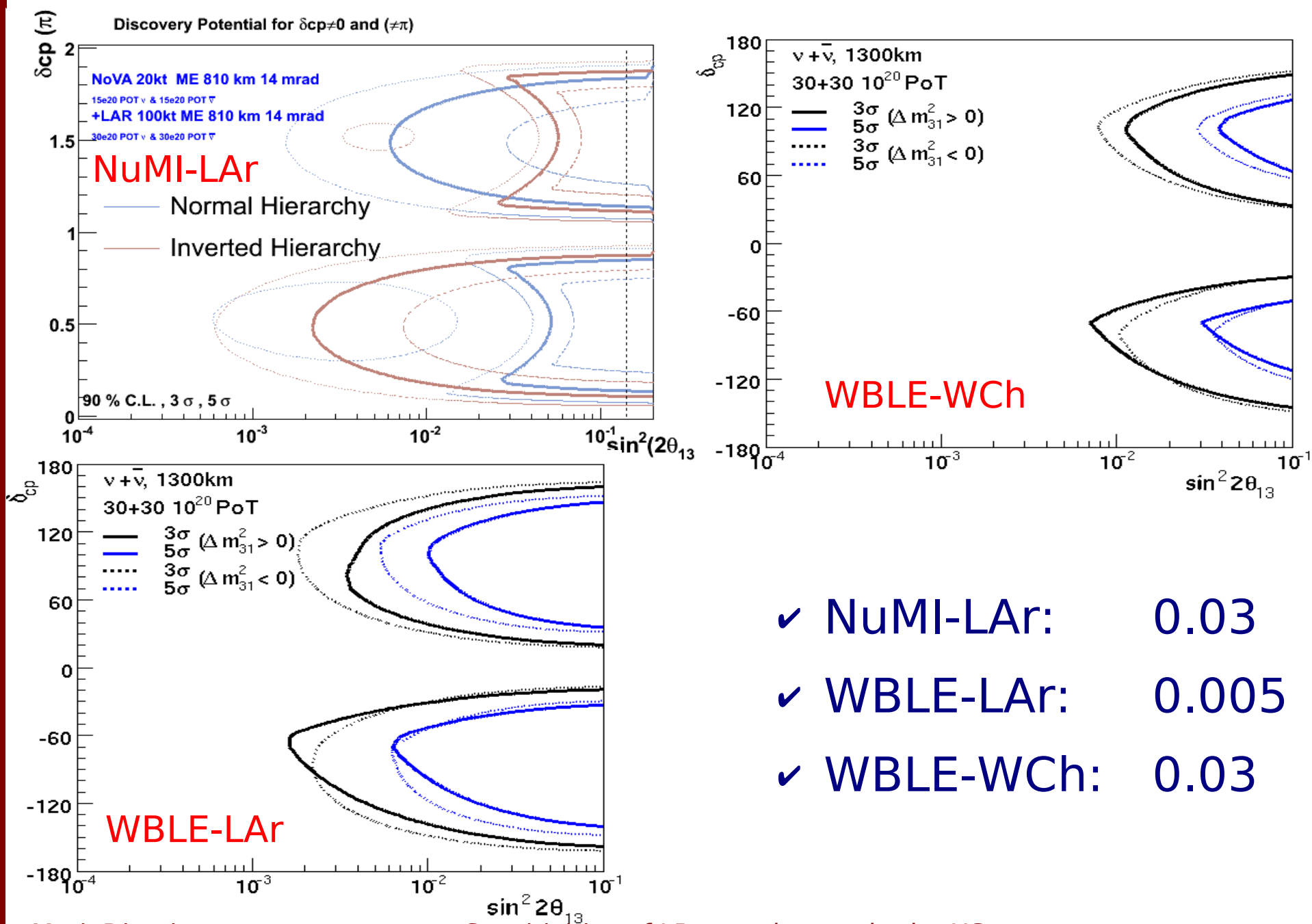
- ✓ Short names for various experiments:
 - NuMI-LAr: 100kt LAr det. in 0.8° off-axis NuMI beam at 810km
 - WBLE-WCh: 300kt Water Cherenkov det. in 120 GeV 0.5° off-axis wide band beam at 1300km
 - WBLE-LAr: same as above but 100kt LAr det.
- ✓ Limits quoted as the value of $\sin^2 2\theta_{13}$ above which the sensitivity is $\geq 3\sigma$ for 50% of δ_{CP} phases for the worst mass hierarchy case.
- ✓ Reminder: fits shown are for 30×10^{20} pot for neutrino and anti-neutrino running each

Sensitivity to $\sin^2 2\theta_{13} \neq 0$



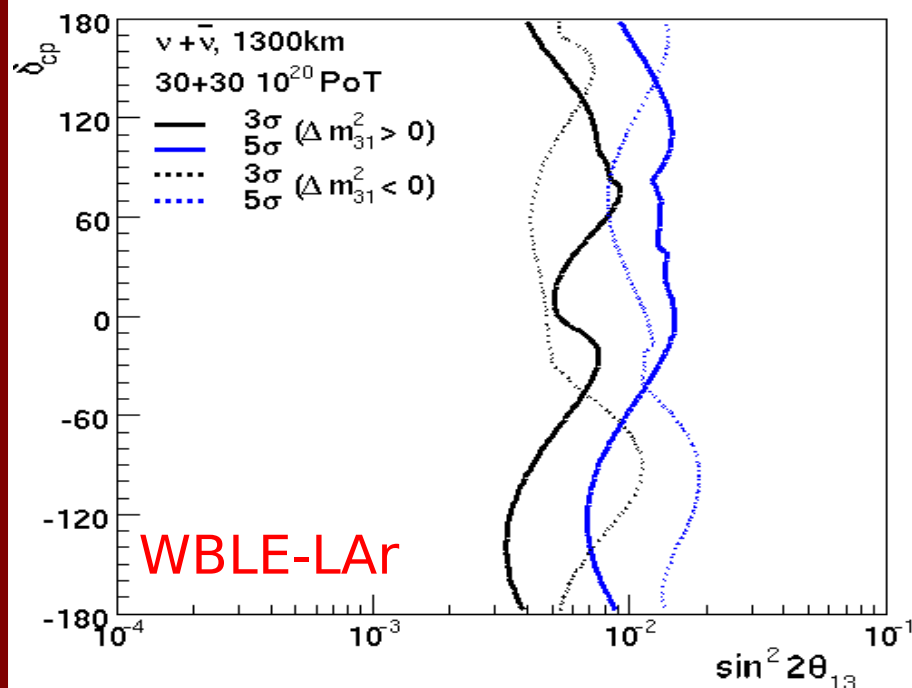
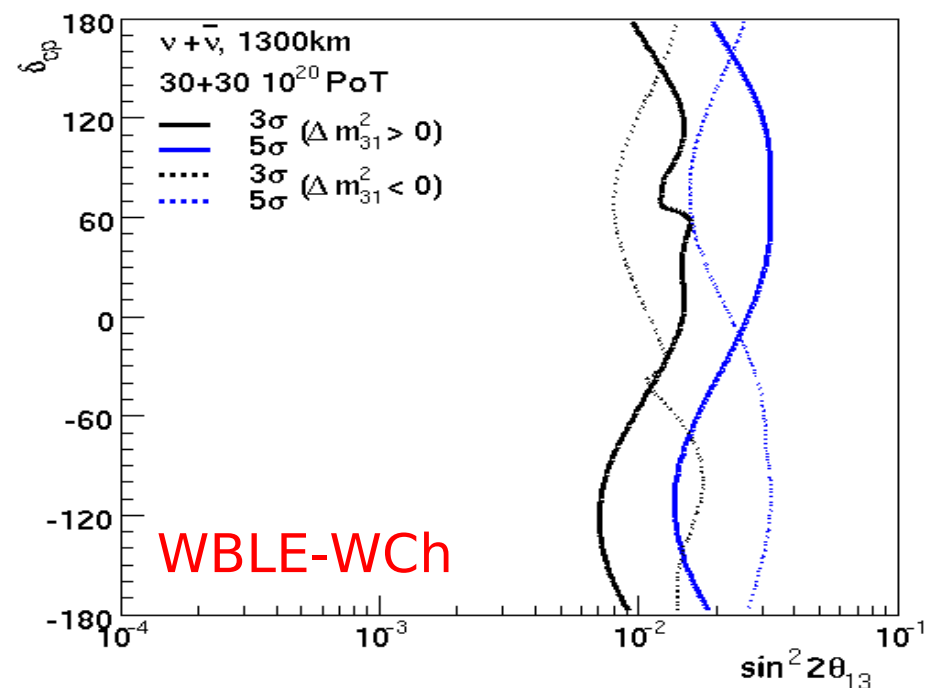
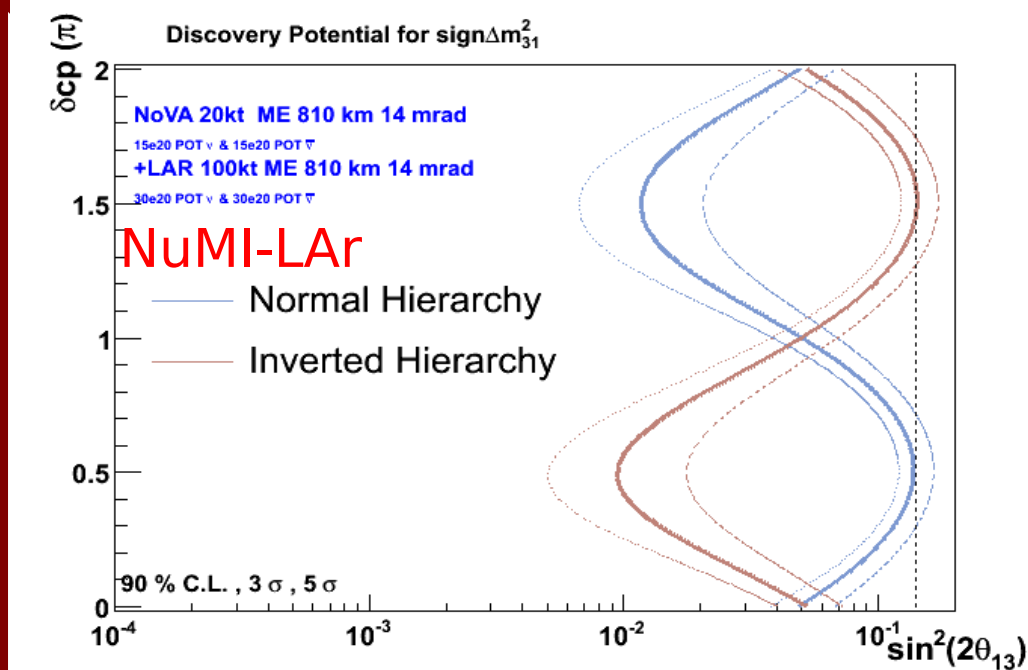
- ✓ NuMI-LAr: 0.002
- ✓ WBLE-LAr: 0.003
- ✓ WBLE-WCh: 0.006

Sensitivity to CP violation



- ✓ NuMI-LAr: 0.03
- ✓ WBLE-LAr: 0.005
- ✓ WBLE-WCh: 0.03

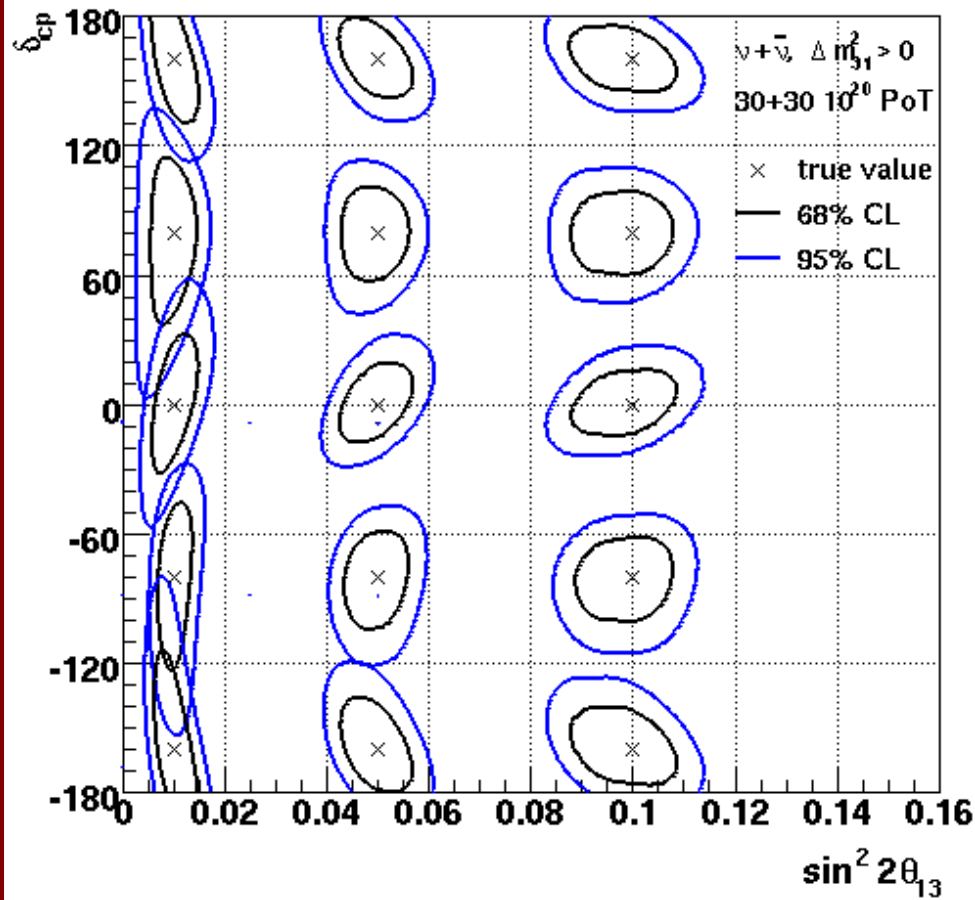
Sensitivity to mass hierarchy



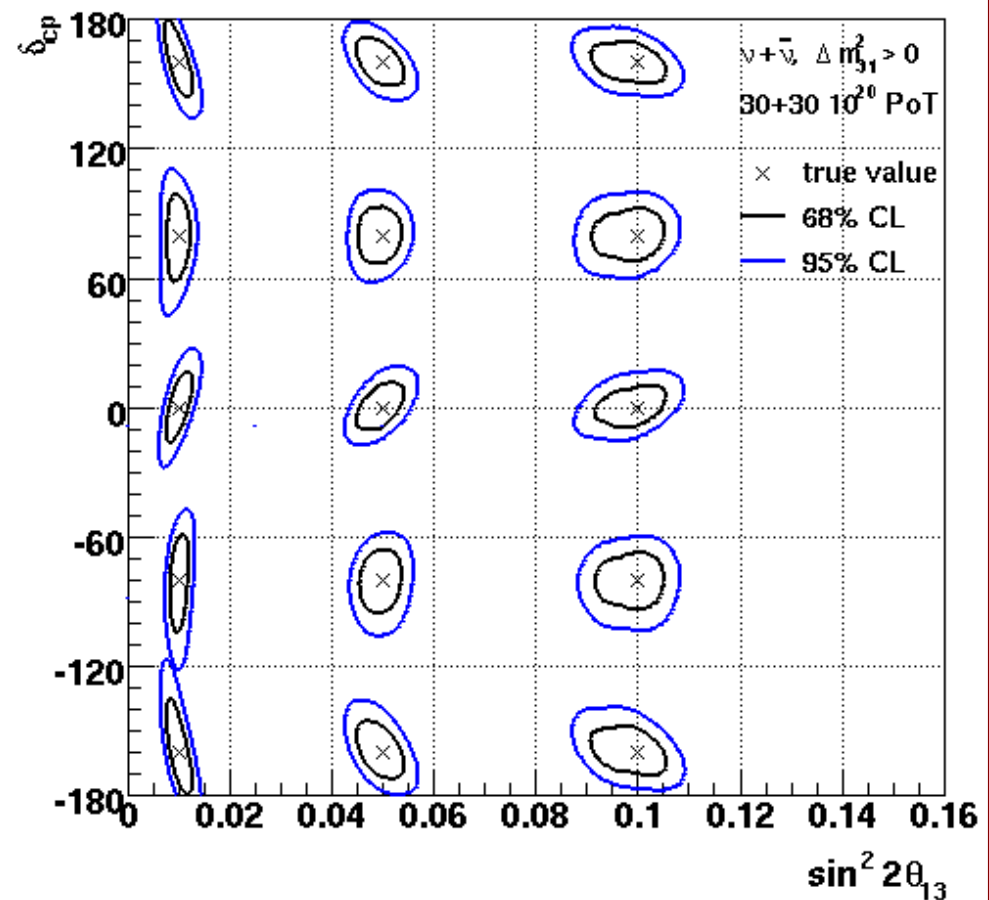
- ✓ NuMI-LAr: 0.05
- ✓ WBLE-LAr: 0.006
- ✓ WBLE-WCh: 0.01

$(\theta_{13}, \delta_{cp})$ measurement

WBLE-WCh



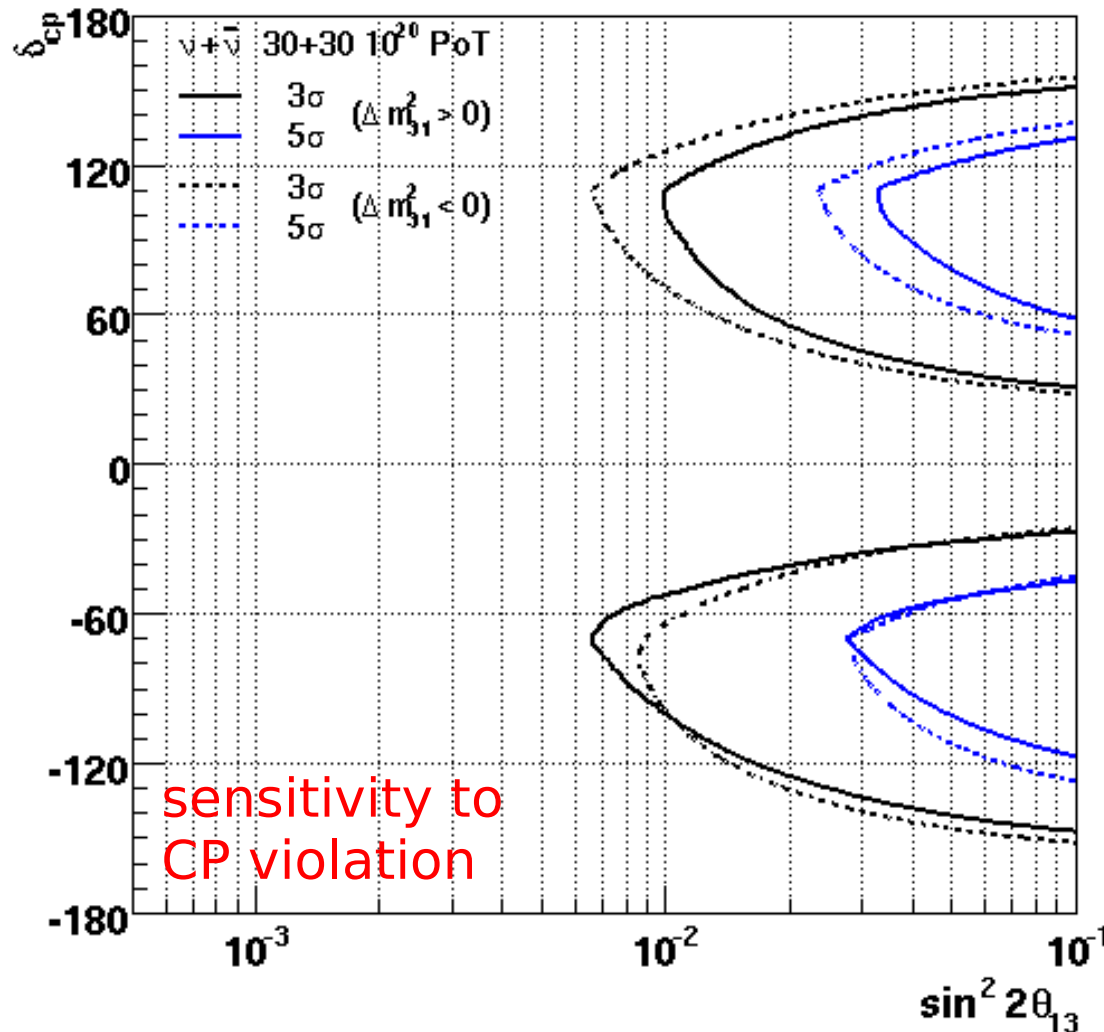
WBLE-LAr



- ✓ Measure $\sin^2 2\theta_{13}$ to 10% (6%) with WBLE-WCh (WBLE-LAr) for $\sin^2 2\theta_{13} > 0.01$ independent of δ_{cp}

Background WBLE-WCh

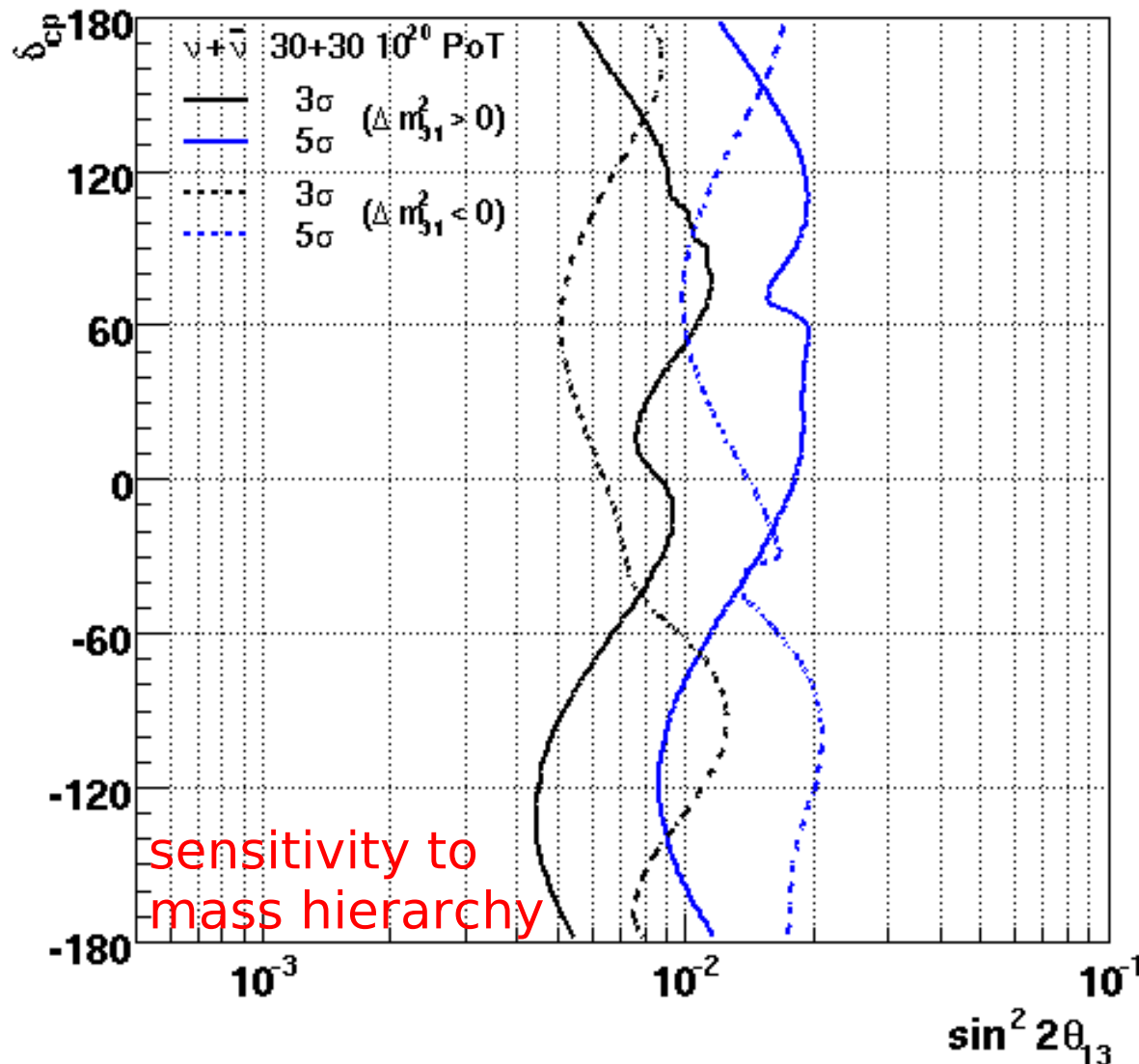
Reduce background uncertainty from 10% to 5%:



- ✓ $\sin^2 2\theta_{13}$:
0.006 \rightarrow 0.005
- ✓ CP violation:
0.03 \rightarrow 0.02
- ✓ $\text{sign}(\Delta m^2_{31})$:
0.01 \rightarrow 0.01

Increased Statistics WBLE-WCh

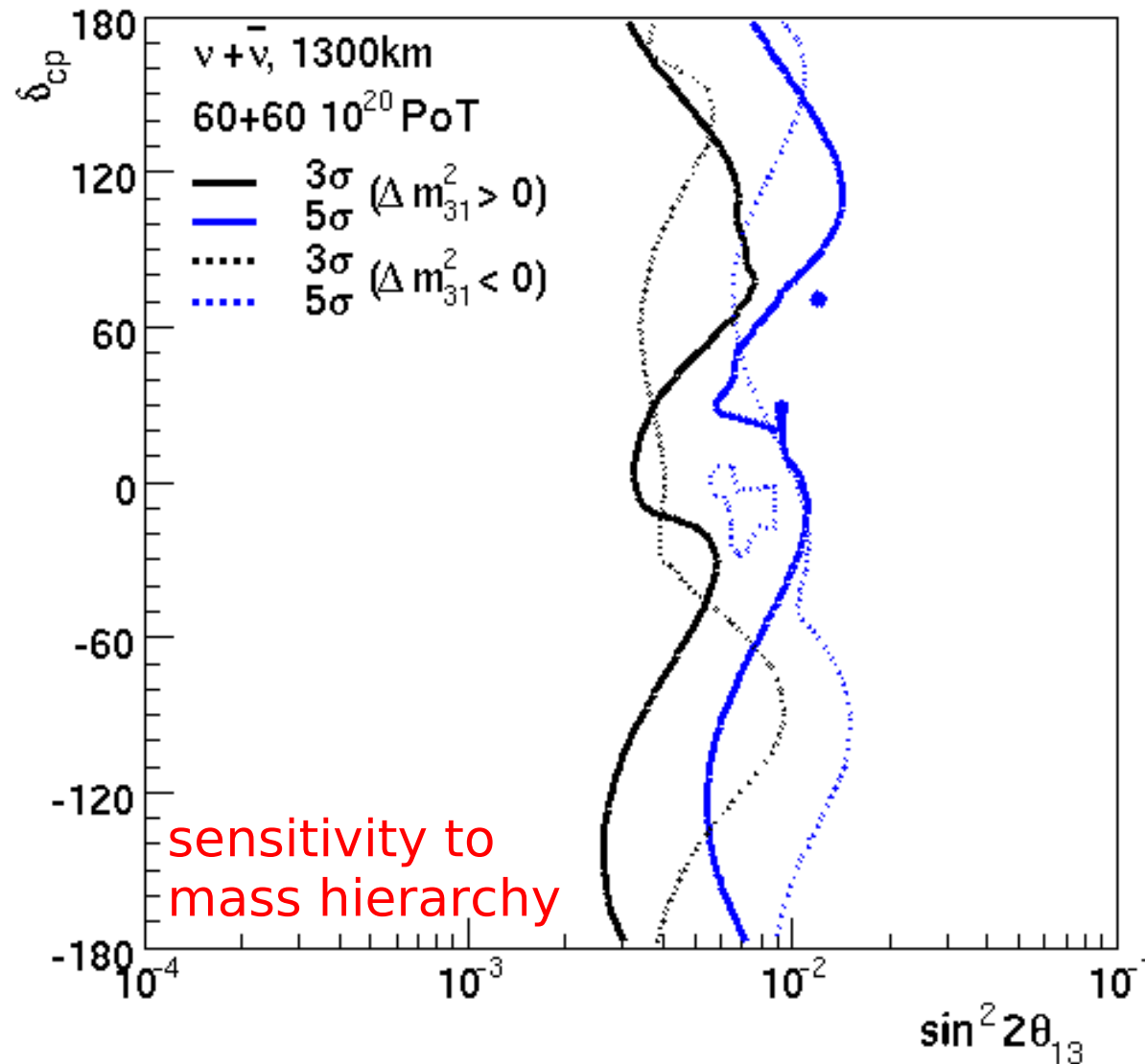
5% background uncertainty & 120 10^{20} pot for $\nu + \bar{\nu}$



- ✓ $\sin^2 2\theta_{13}$:
0.006 \rightarrow 0.004
- ✓ CP violation:
0.03 \rightarrow 0.01
- ✓ $\text{sign}(\Delta m^2_{31})$:
0.01 \rightarrow 0.008

Increased statistics WBLE-LAr

60 10^{20} pot for ν and $\bar{\nu}$ each



$\sin^2 2\theta_{13}$:

$0.003 \rightarrow 0.002$

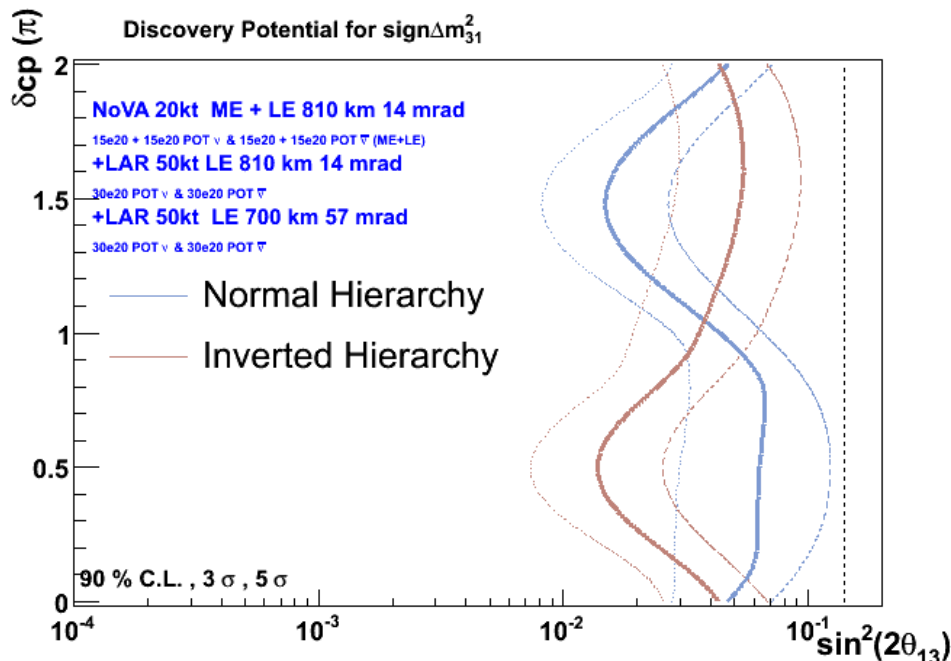
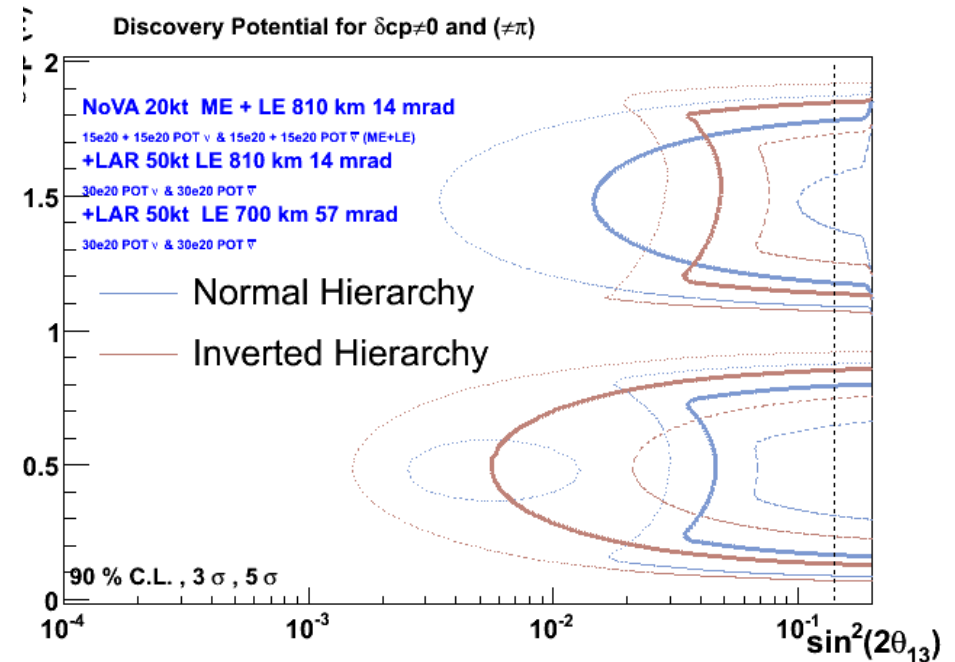
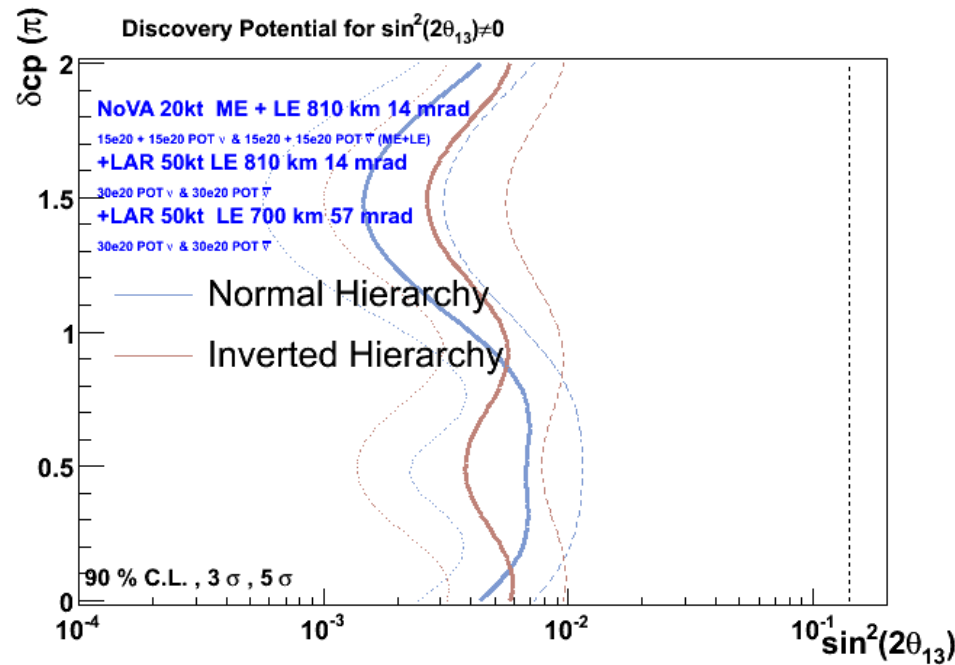
CP violation

$0.005 \rightarrow 0.003$

$\text{sign}(\Delta m^2_{31})$:

$0.006 \rightarrow 0.004$

2nd off-axis detector



50kt LAr 0.8° &
50kt LAr 3° off axis:

✓ $\sin^2 2\theta_{13}$: 0.005

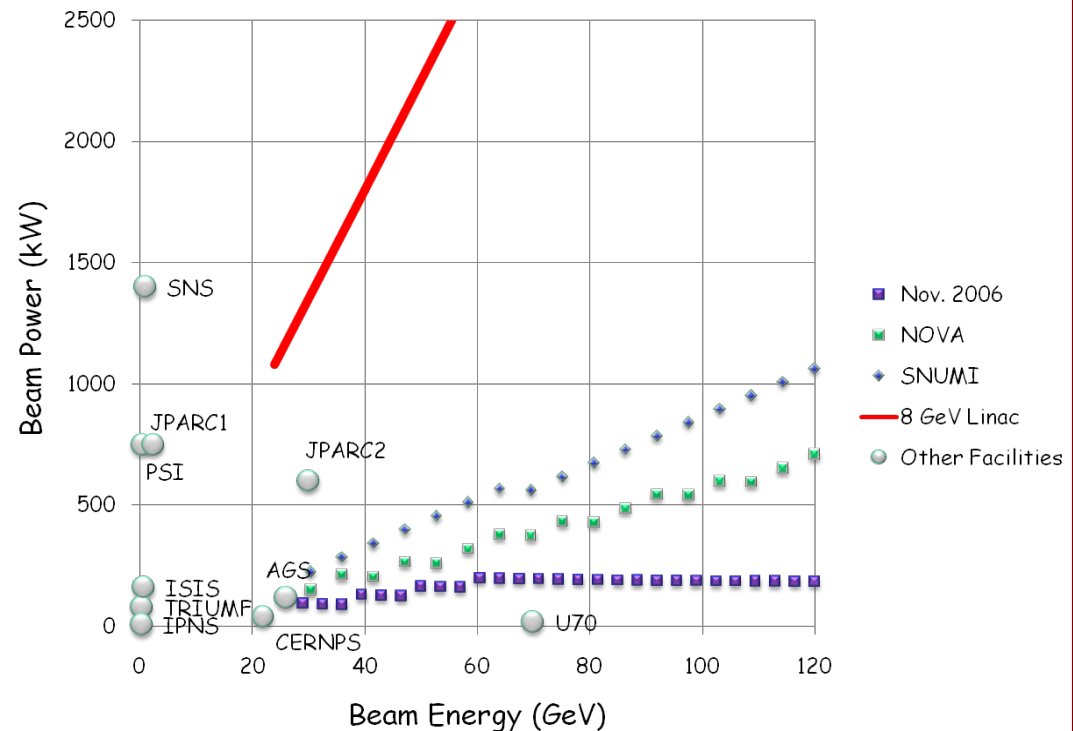
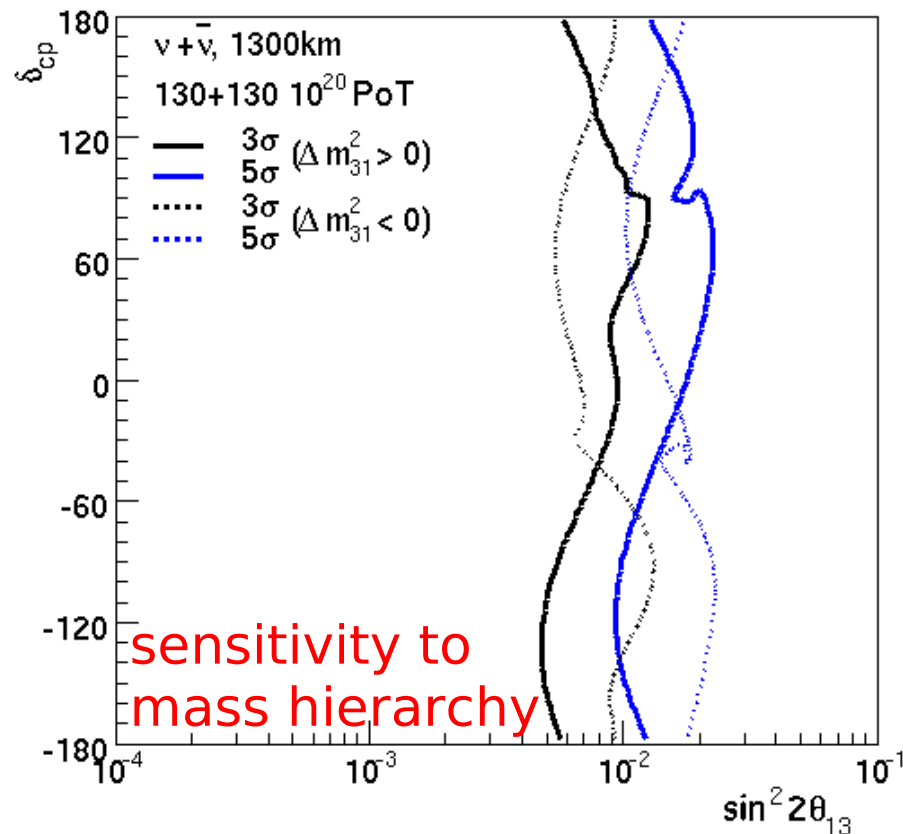
✓ δ_{CP} : 0.04

✓ $\text{sign}(\Delta m_{31}^2)$: 0.04

Project X

- ✓ From Y.K. Kim's talk at HEPAP meeting July 07: 2.6MW at 60GeV

<http://www.science.doe.gov/hep/HEPAPJuly2007Agenda.htm>



✓ $\sin^2 2\theta_{13}$: 0.003

✓ δ : 0.01

✓ $\text{sign}(\Delta m_{31}^2)$: 0.007

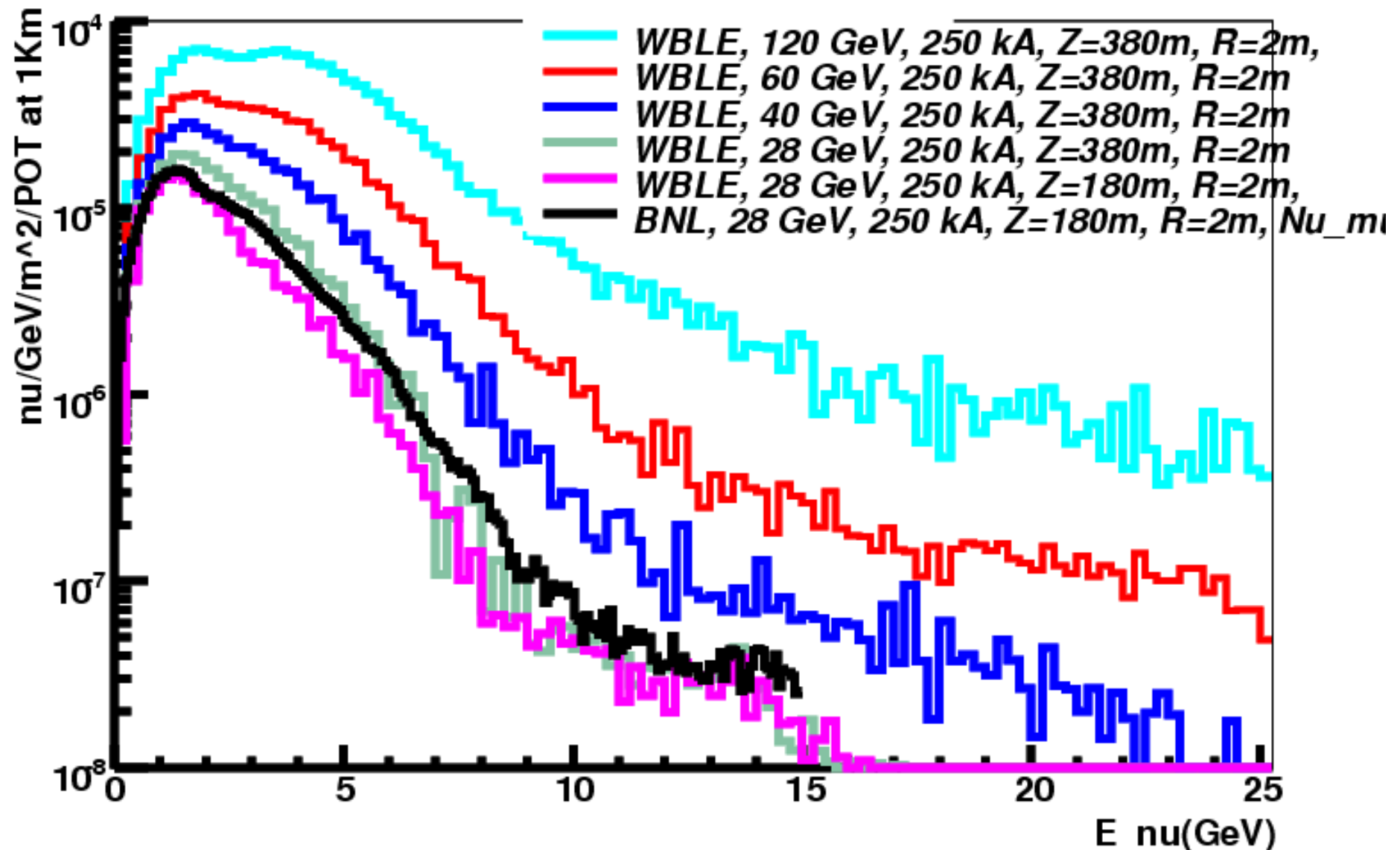
Conclusions

still to write...

- ✓ More information and supporting documents available at :
<http://nwg.phy.bnl.gov/~diwan/nwg/fnal-bnl/>

Backup Slides

Wide band beam



WCh detector

- ✓ put some plots from Chiaki/Fanny

Background WBLW-WCh

- ✓ put plots with x2 background reduction

Other Off-axis combinations